A Decade of Radioactive Beams at ISAC Reflecting on ISAC's 10 Year History First Beam (<sup>36m</sup>K) – November 1998

> John M. D'Auria Simon Fraser University

- ➤ What is ISAC?
- > What is its place in the World?
- ➢ How Did It Start?
- Science in the First 10 Years?
- > What Does the Future Hold?
- Concluding Remarks?

### What is ISAC?





Latest generation of an on-line isotope separator (ISOL)

Elemental selectivity based upon the Target and Ion Source; Mass selectivity based upon Mass Analyzer

**Chemistry and Physics** 

Coupling of an ISOL with a post-accelerator to produce energetic RB (essentially the first in the world)



### **ISAC @ TRIUMF**



Target material for beam production includes SiC, CaO, TiC, U (UO and UC license) Ion sources: surface, laser, FEBIAD, ECR (test)

#### **ISAC:**

Highest power for On-Line facilities, presently up to 100µA @ 500MeV DC proton

#### ISAC has 3 exper. areas:

- Low energy (< 60keV)
- ISAC I (150 keV/u –1.8 MeV/u)
- ISAC II (up to 16MeV/u, presently being upgraded)
- Suite of experimental stations:
- TRINAT, Beta-NMR, 8pi, tapestation, TITAN, Co-linear laser spec, polarised beam line, etc
- DRAGON, TUDA, TACTIC, GPS,
- TIGRESS ( 8/12),
- EMMA (2011),
- HERACLES

Some Beam Intensities @ Yield Station

<sup>8</sup> Li	(Ta)	8 x 10 <sup>8</sup> pps
<sup>11</sup> Li	(Ta)	4 x 10 <sup>4</sup> pps
<sup>21</sup> Na	(SiC)	9.9 x 10 <sup>9</sup> pps
<sup>26g</sup> Al	(SiC)	1 x 10 <sup>10</sup> pps
<sup>74</sup> Rb	(Nb)	1.3 x 10 <sup>4</sup> pps
<sup>79</sup> Rb	(Nb)	4.6 x 10 <sup>9</sup> pps
<sup>160</sup> Yb	(Ta)	8.4 x 10 <sup>9</sup> pps

### World of Radioactive Beams Facilities



ISAC presently produces highest intensity RB in world (for selected beams)

How did it start? (Pre-ISAC Highlights)

- 1951 First on-line ISOL (Niels Bohr Inst.)
- 1969–present ISOLDE (CERN)
- '60-80's- Various ISOL facilities (OSIRIS, ISOCEL, TRISTAN, UNISOR, ...)
- ~1970's John Warren suggests an ISOL on BL1;
- ~1980's Proposal to move TRISTAN to BL4A
- 1985 Proposal to TRIUMF BOM

Conferences at Mount Gabriel and Parksville

• 1986–1999 TISOL (R&D and Production RB facility)

(Mini Workshops at Lake Harrison (WEISS), and Lake Louise

- 1995 ISAC Funding (and Construction starts)
- 1998 First beam (<sup>38m</sup>K) to TRINAT

## 1951: First ISOL beams at NBI O. K Hansen and K.O. Nielsen



Figure 3. The Copenhagen isotope separator.

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TRI-85-1

PROCEEDINGS OF THE ACCELERATED RADIOACTIVE BEAMS WORKSHOP

> PARKSVILLE, CANADA September 5-7, 1985

Editors: L. Buchmann, University of Toronto J.M. D'Auria, Simon Fraser University

- Organized by: C.A. Barnes, California Institute of Technology L. Buchmann, University of Toronto J.M. D'Auria, Simon Fraser University C. Rolfs, Universität Münster
- Sponsored by: Natural Sciences and Engineering Research Council of Canada National Science Foundation (U.S.A.) Deutsche Forschungsgemeinschaft (West Germany) TRILUMF Simon Fraser University



#### **Key Players:**

- R. Azuma-UToronto J.K.P. Lee McGill J. Crawford, McGill R. Moore, McGill C. Rolfs, Munster J. King, UToronto L. Buchmann, UT
- J. D'Auria, SFU

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#### ISAC Isotope Separator/Accelerator

ISAC will accelerate radioactive isotopes to high velocities, a capability which will allow scientists to replicate reactions which occur in stars in the distant universe, and to study nuclear structure, the behavior of unusual atomic nuclei, condensed matter physics and life sciences projects. This is recognized worldwide as leading edge research and will place Canada back in the forefront of nuclear physics. The world physics community wanted ISAC, and Canada was in a unique position to develop it because TRIUMF can use the high power proton beam from the existing TRIUMF cyclotron to produce the copious beams of exotic, short-lived radioisotopes needed for the ISAC facility.

- 1 ISAC ground breaking, April 1996
- 2 Installing first shielding in the target hall
- 3 Aerial view of TRIUMF
- 4 ISAC building from the southeast
- 5 Early construction phase of the target hall
- 6 Beam line 2A transports the proton beam from the TRIUMF cyclotron to the target in the new ISAC facility
- 7 Installing the Radio Frequency Quadrupole tank
- 8 Tunnel construction from the cyclotron to the ISAC Facility
- 9 Completed ISAC building, February 1998



# ISAC Technical Highlights

(First 10 Years since First Beam)

# **Major Milestone**

(ignoring infrastructure construction)

- 1998<sup>\*\*\*</sup> First RB beam ( $^{38m}$ K) to TRINAT;
- 1998/9 LEBT from OLIS to RFQ; RFQ full power
- 2000 First physics (<sup>74</sup>Rb lifetime with high precision)

(R. Boyd experiment??)

- 2001 Accelerated beam of <sup>21</sup>Na to TUDA and DRAGON
- 2003 TITAN and TIGRESS Funded; ISAC II bldg. opened
- 2004 Laser ion source used for exp.; CSB tested;<sup>26g</sup>Al beam found; high power target developed
- 2005 <sup>4</sup>He beam in single ISAC II module
- $2006 {}^{40}$ Ca beam accelerated to 5.5 MeV/u in ISAC II
- 2007 Energetic <sup>11</sup>Li beam delivered to Maya Experiment in ISAC II

\*\*\* Remember funding in 1995 (built on budget and on schedule)

### **Challenges and Achievements**

- RFQ LINAC (Room temperature, low frequency, 150 keV 1.8 MeV/u)
- High power targets (Handling up to 5 kWatts)
- Thin ISOL targets to allow fast release of very short isotopes
- Selection of ion sources (surface, FEBIAD, LASER, ECR?)
- Unique shielding and remote handling (using cranes not robots)
- Professional Operations (as compared to older ISOL facilities)
- Proper Remote Control Systems (EPICS based)
- Excellent collection of world class, experimental facilities
- Fixed beam scheduling (Unlike HRIBF, NSCL, others?)
- OLIS (almost forgotten initially)

# Scientific Highlights (First 10 Years)

# Science at ISAC The First 10 Years

- Nuclear Astrophysics
  - DRAGON
  - TUDA
  - DAT
- Fundamental Physics
  - TRINAT
  - Precision Lifetime Measurements
- Condensed Matter Physics
  - Beta NMR and Polarized <sup>8</sup>Li
  - low field beta-NMR and zero field beta-NQR
- Nuclear Structure
  - The  $8\pi$  (and peripheral detector arrays)
  - TIGRESS
  - Halo Nuclei Experiments
  - MAYA experiment
- Nuclear Masses
  - TITAN (CPT??)

# **Nuclear Astrophysics**

DRAGON TUDA DTM



Thanks to Witek Nazarewicz, U. Tennessee/ Greg Hackman

### ISAC - Poised to Measure Nuclear Astrophysics



# DRAGON

C. Ruiz

#### Detector of Recoils and Gammas of Nuclear Reactions

Direct measurement of radiative capture reactions at (explosive) stellar energies, in inverse kinematics

- MEME design
- Windowless recirculating target: H<sub>2</sub> or He
- LN2 zeolite cleaning trap
- 1-10 mbar
- 30-element BGO gamma array;40-70% efficiency

Combination of RB + RFQ + DRAGON is unique in world



# DRAGON Program Summary

- 5 radiative capture reactions measured directly using RIBs in last 20 years (world).
  4 in last 8 years: 3 using DRAGON
- Textbook RIB measurement: <sup>21</sup>Na(p,γ)<sup>22</sup>Mg for <sup>22</sup>Na synthesis in O-Ne novae All contributing resonances measured to high precision: reaction rate completely determined experimentally; *most intense* (~10<sup>9</sup>/s) <sup>21</sup>Na RIB in world.
- New cascade transition in <sup>12</sup>C(α,γ)<sup>16</sup>O
  DRAGON sensitive to transitions unobserved in other experiments
- First inverse kinematics measurement of <sup>26g</sup>Al(p,γ)<sup>27</sup>Si, highest intensity (~10<sup>9</sup>/s) <sup>26g</sup>Al beam in world
- Direct measurement of <sup>40</sup>Ca(α,γ)<sup>44</sup>Ti in inverse kinematics for supernova <sup>44</sup>Ti production, *best measurement* so far
- > *First direct measurement* of  ${}^{23}Mg(p,\gamma){}^{24}AI$  reaction using *most intense*  ${}^{23}Mg$  RIB
- > Production of Target of <sup>22</sup>Na to perform ( $p,\gamma$ ) for <sup>22</sup>Na synthesis in O-Ne novae

# The ${}^{21}Na(p,\gamma){}^{22}Mg$ reaction





A.F. Iyudin et al, Astron. Astrophys. 300, 422 (1995)

- 21Na(p, $\gamma$ )<sup>22</sup>Mg bypass for <sup>21</sup>Na( $\beta^+$ )<sup>21</sup>Ne(p, $\gamma$ )<sup>22</sup>Na route
- Stronger <sup>21</sup>Na(p,γ)<sup>22</sup>Mg means <sup>22</sup>Na destroyed quicker
- Ejected <sup>22</sup>Na abundance strongly dependent
- Predicted abundance in O-Ne novae ~6 x 10<sup>-9</sup> solar masses
- <sup>22</sup>Na: prospect of observation from single nearby (< 1kpc)\* nova</li>
- No observation by COMPTEL (Nova Her 1991 ~3kpc, Nova Cyg 1992 ~2kpc)
- INTEGRAL slightly more sensitive
- When observation comes, need accurate <sup>22</sup>Na ejected yield estimations



#### **APS DNP Fall Meeting**

#### RESULTS

Phys. Rev. Lett. 90, 16 (2003) Phys. Rev. C 69, 065803 (2004)



### Level structure <sup>22</sup>Mg



## Conclusions on ${}^{21}Na(p,\gamma){}^{22}Mg$ direct measurement

- <sup>21</sup>Na beam 'easy'; prolific from spallation in SiC, readily ionizeable in rhenium
- Pure beam (no isobars)
- Primary beam radioactive and intense: normalization via elastic scattering and beta monitor sufficient
- Dominant experimental background: primary beam decay shielding & high thresholds required
- High intensity makes thick target excitation functions possible in reasonable time
- Rate determined to +/- 20% precision
- 'Classic' DRAGON experiment

Spinoff: DRAGON measurement of  $E_R$ =206 keV <sup>21</sup>Na(p, $\gamma$ )<sup>22</sup>Mg resonance revealed a 6 keV discrepancy in the reaction Q-value, traced to the mass of <sup>22</sup>Mg

This has since been corroborated by re-evaluation: J. Hardy, Phys. Rev. Lett. 91 (2003)

## <sup>40</sup>Ca(α,γ)<sup>44</sup>Ti at DRAGON <sup>44</sup>Ti: youngest indicator of nucleosynthesis



60-year half-life, observed in Cas A, SN1987A

<sup>44</sup>Ti produced by <sup>40</sup>Ca( $\alpha$ ,γ)<sup>44</sup>Ti in 'alpha-rich freeze-out'

Large discrepancy between previous experiments



#### **DRAGON** measurement



**40% increase** of <sup>44</sup>Ti yield compared to empirical model (Rauscher *et al.* 2000)

DRAGON measured uncertainty of the reaction rate = +/-3% in <sup>44</sup>Ti yield

<sup>44</sup>Ti yield and a <sup>44</sup>Ti/ <sup>56</sup>Ni ratio agree better with observations in Cas A and SN1987A. [cf. Nassar *et al.*, *Phys. Rev. Lett.* 96 (2006)]

Measurement of the  ${}^{40}Ca(\alpha, \gamma){}^{44}Ti$  reaction relevant for supernova nucleosynthesis

C. Vockenhuber,<sup>1,\*</sup> C.O. Ouellet,<sup>2</sup> L.-S. The,<sup>3</sup> L. Buchmann,<sup>1</sup> J. Caggiano,<sup>1</sup> A.A. Chen,<sup>2</sup> H. Crawford,<sup>1</sup> J.M. D'Auria,<sup>4</sup> B. Davids,<sup>1</sup> L. Fogarty,<sup>1</sup> D. Frekers,<sup>5</sup> A. Hussein,<sup>6</sup> D.A. Hutcheon,<sup>1</sup> W. Kutschera,<sup>7</sup> A.M. Laird,<sup>8</sup> R. Lewis,<sup>8</sup> E. O'Connor,<sup>1</sup> D. Ottewell,<sup>1</sup> M. Paul,<sup>9</sup> M.M. Pavan,<sup>1</sup> J. Pearson,<sup>2</sup> C. Ruiz,<sup>1</sup> G. Ruprecht,<sup>1</sup> M. Trinczek,<sup>1</sup> B. Wales,<sup>2</sup> and A. Wallner<sup>7</sup>

### TUDA/TACTIC: charged particle reactions

- TUDA (TRIUMF-UK Detector Array): direct charged particle reactions or indirect techniques (resonant elastic scattering)
- 5 RIB experiments:
  - <sup>21</sup>Na(p,p)<sup>21</sup>Na resonant elastic scattering
  - <sup>20</sup>Na(p,p)<sup>20</sup>Na resonant elastic scattering
  - <sup>18</sup>F(p, $\alpha$ )<sup>15</sup>O direct , <sup>18</sup>F(p,p)<sup>18</sup>F res.
  - ${}^{21}$ Na(p, $\alpha$ )<sup>18</sup>Ne indirect (XRB)
  - <sup>7</sup>Li(<sup>8</sup>Li,<sup>7</sup>Li)<sup>8</sup>Li ANC indirect (S<sub>17</sub>(0) test of IMS)



#### TACTIC

# (TRIUMF Annular Chamber for Tracking and Identification of Charged Particles

 $^{8}\text{Li}(\alpha,n)^{11}\text{B}$  for BBN and r-process

 $^7\text{Li}(^3\text{He},\alpha)^6\text{Li}$  for BBN

First operation 2008

Group leader: L. Buchmann

### Indirect techniques: Doppler Shift Lifetime Facility

S. Mythili, B. Davids et al. Phys. Rev. C

R. Kanungo et al. Phys. Rev. C. 74

77 (2008)

(2007)

- Populate excited states of interest to measure lifetimes using Doppler Shift Attenuation Method (at ISAC I or ISAC II)
- Use of TIGRESS detectors
- Custom chamber and implanted targets
- ${}^{15}O(\alpha,\gamma){}^{19}Ne$  and  ${}^{14}N(p,\gamma){}^{15}O$  studied indirectly in this way







Group Leader: Barry Davids

# Fundamental Physics Symmetries

## TRINAT Superallowed Beta Decay

# Physics justification:

Nuclear Beta Decay



## TRINAT TRIUMF's Neutral Atom Trap

- Isotope/Isomer selective
- $\bullet$  Evade 1000x untrapped atom background by  $\rightarrow$  2nd MOT
- 75% transfer (must avoid backgrounds!); 10<sup>-3</sup> capture
- 0.7 mm cloud for  $\beta$ -Ar<sup>+</sup>  $\rightarrow \nu$  momentum  $\rightarrow \beta$ - $\nu$  correlation





### TRIUMF Neutral Atom Trap

The pressure of laser light traps atoms of unstable isotopes in a 1mm-sized cloud.

The final atoms after decay have small kinetic energies (~100 eV) but freely escape the trap so we can measure their momentum. In 'beta decay' we measure the electron and atom momentum and deduce the (otherwise invisible) neutrino momentum.

SAC ion beam



# **TRINAT Results and Impact**

TRIUMF Neutral atom trap:  $\beta$  decay correlations

- Upgrades in progress, goals 5-10 X better:
  - <sup>38m</sup>K  $\beta$ - $\nu$  correlation: Gorelov PRL 2005, best general limits on scalars coupling to 1st generation • <sup>37</sup>K :  $\nu$  spin asymmetry Melconian PLB 2007 B<sub> $\nu$ </sub>/B<sup>SM</sup><sub> $\nu$ </sub> = 0.982 ± 0.026 ± 0.017

 Singles recoil spin asymmetry <sup>80</sup>Rb Pitcairn PRC 2009 A<sub>recoil</sub>=0.015±0.029±0.019 Complementary constraints on tensor interactions



590

goals: scalar, tensor interactions from SUSY produce  $\sim$  0.001 effects (Profumo, Ramsey-Musolf, Tulin PRD 2007) Searching for exotic massive particles in 2-body decay

John Behr

### Precision Lifetime Measurements at ISAC Beta-decay tape station



4  $\pi$  gas-proportional  $\beta$  counter setup





G. Ball

Superallowed Beta Decay/Precision Lifetime Measurements

### The Cabibbo-Kobayashi-Maskawa (CKM) matrix

 The CKM matrix plays a central role in the Standard Model describes the mixing of different quark generations:
 weak interaction eigenstates ≠ quark mass eigenstates







$$\left| d' \right\rangle = V_{ud} \left| d \right\rangle + V_{us} \left| s \right\rangle + V_{ub} \left| b \right\rangle$$

- In the Standard Model the CKM describes
- a unitary transformation.

$$V_{ud}^2 + V_{us}^2 + V_{ub}^2 = 1$$

 $\begin{pmatrix} d'\\ s'\\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d\\ s\\ b \end{pmatrix}$ 

The first row of the CKM matrix provides the most demanding experimental test of this unitarity condition.


For the special case of  $0^+ \rightarrow 0^+$  (pure Fermi)  $\beta$  decays between isobaric analogue states (superallowed) the matrix element is that of an isospin ladder operator:

 $|M_{fi}|^2 = (T - T_Z)(T + T_Z + 1) = 2$  (for T=1)

Strategy: Measure superallowed ft-values, deduce  $G_V$  and  $V_{ud}$ :

Vector coupling 
$$\longrightarrow G_V^2 = \frac{K}{2 \text{ ft}}$$
  $|V_{ud}| = G_V / G_F \leftarrow \text{Fermi coupling constant}$ 

## Superallowed Fermi β Decays





# **Nuclear Structure**

- Halo Nuclei Studies



#### The $8\pi$ at TRIUMF

- 20 25% HPGe w/ Suppressors
- Endless loop moving tape system (LSU)
- Programmable states (beam on/off, tape move, trigger veto)
- Inner plastic scintillator array,
- Options: Si(Li), BaF<sub>2</sub>
- For <sup>19</sup>Ne:
  - Plastics for lifetime
  - HPGe for branching ratio







#### TRIUMF-ISAC Gamma-Ray Escape-Suppressed Spectrometer (TIGRESS):







Four ~40% n-type HPGe crystals close-packed in a four-leaf clover geometry.
32-fold segmentation of the outer contacts provide position resolution.
Even more powerful when combined with auxiliary detectors.
University of Guelph, McMaster University, Université de Montréal, University of Toronto, Université Laval, Simon Fraser University, and TRIUMF



### TIGRESS Ge Gamma Array and SHARC Silicon barrel





















## Halo Nuclei (radius>> $r_0 A^{1/3}$ )

#### **Nuclear halos**



#### Selected Recent Publications on Halo Nuclei Experiments at TRIUMF-ISAC

- <sup>11</sup>Li β-n-DSAM : F. Sarazin *et al.*, Phys. Rev. C 70, 031302R (2004), and C. Mattoon et al., PRC 80, 034318 (2009) [*CSM*]
- □ <sup>11</sup>Li β-n: Y. Hirayama *et al.*, Phys. Lett. B611, 239 (2005) [Osaka]
- □ <sup>11</sup>Li charge radius: R. Sanchez *et al.*, PRL 96, 033002 (2006) [*GSI, U. Tübignen*]
- □ <sup>11</sup>Li β-charged particle: R. Raabe *et al.*, PRL 101, 212501 (2008) [*K.U.Leuven*]
- <sup>11</sup>Li two-neutron transfer (p,t): Tanihata, Savajols *et a*l, PRL 100, 192502 (2008) [GANIL,Osaka]
- □ <sup>9,11</sup>Li mass: M. Smith *et al.*, PRL 101, 202501 (2008)
- □ <sup>8</sup>He mass: V. Rykov *et al.*, PRL 101, 202301 (2008)
- □ <sup>11</sup>Be mass: R. Ringle *et al.*, PLB 675, 170 (2009)

## The Maya Experiments

- 1. The p(<sup>11</sup>Li,<sup>9</sup>Li)t reaction at 3 MeV/u
  - First study at ISAC 2
  - Large international group
  - time projection chamber from GANIL
  - Studying correlations between halo n
  - 2500-5000 pps of <sup>11</sup>Li on target
  - transitions to gs and 1st ex state





FIG. 1. (Color online) A schematic drawing of the MAYA detector. The <sup>11</sup>Li beam is incident from the left, and ionization electrons drift down to the anode wires. Projections of the charged particles trajectories are recorded on the honeycomb segmented cathode. Escaping particles hit the backside Si + CsI wall.

- 2. Mass of Li from the p,t reaction
  - mass derived from the Q value of the p,t reaction
  - Q value of 8.119(22) MeV
  - S<sub>2n</sub> = 363(22) MeV consistent
     with results from TITAN and MISTRAL

# **Nuclear Masses**

#### **TITAN mass measurement system**

Jens Dilling



### Strange form of matter: Halo nuclei an 'old' phenomena, but new methods at ISAC



11

## Halo mass measurements: <sup>11</sup>Li



- TITAN mass measurement of <sup>6,7,8,9,11</sup>Li
- Improved precision, S<sub>2n</sub> improved by factor 7
- Shortest-lived isotope (T<sub>1/2</sub>=8.8ms) for Penning trap mass measurement!
- Final analysis δm = 650 eV
- M. Smith et al PRL 101, 202501 (2008)
- Re-evaluation of radius, and comparison to theory



## TITAN Masses – K Isotopes

• <sup>44</sup>K<sup>4+</sup>: First on-line mass measurement using charged bred ions from the EBIT

• <sup>47-50</sup>K<sup>1+</sup> and <sup>49,50</sup>Ca<sup>1+</sup>: masses improved by factor of up to 100

•  ${}^{48}K^{1+}$  and  ${}^{49}K^{1+}$ : deviations of 6 and 10  $\sigma$  from AME03



## **Condensed Matter Physics**

- Coupling of use of radioactive beams (nuclear) for Material science
- (Otto Hauser brought Rob Kiefl to the table at Traps Workshop)
- Presently uses <sup>8</sup>Li (and <sup>9</sup>Li) beams, 10<sup>9</sup>/s (could use other beams also)
- Remember T<sub>1/2</sub>(<sup>8</sup>Li) = 838 ms !!!
- 30-60 keV,
- Polarized !!!
- Sample on HV platform to tune implantation depth
- Adjust temperature, B fields, RF
- Now there are 2 experimental stations
  - High field beta-NMR
  - Low field beta-NMR and zero field beta-NQR
- Detect <sup>8</sup>Li beta asymmetry
- Ideal for senstive probing of thin films
- Connects with  $\mu$ SR studies (very large user group)
- Brings many users to ISAC also

## $\beta$ -NMR Spectrometers at ISAC



Electrostatic deceleration is used to control the depth of the implanted ions (2-500nm)

#### β-NMR Spectrometers at ISAC



Polarizer



## Thin Film Ag on Nb Superconductor

10

12

Frequency (kHz)

14

•Thin Ag film on Nb surface Exhibits "Proximity Effect" Superconductivity



•G. D. Morris, W.A. MacFarlane, R. Keifl





Fig. 1: Left: Resonance from <sup>8</sup>Li in Ag above and below the critical temperature T<sub>cNS</sub> of the Ag(40nm)Nb(300nm) bilayer. Right: Resonance peak value versus temperature showing the diamagnetic shift below  $T_{cNS}$  in Ag.

Temperature (K)

16

#### Measurement of the quadrupole moment of <sup>9</sup>Li and <sup>11</sup>Li



# Future for ISAC

# EMMA, the ElectroMagnetic Mass Analyser



1 m

Recoil Mass Spectrometer for ISAC-II Project Leader: Barry Davids Length = 9 m; 1<sup>st</sup> order mass resolving power = 500 Solid angle:  $\pm 4^{\circ}$  by  $\pm 4^{\circ} = 20$  msr M/q acceptance =  $\pm 4\%$ ; Energy acceptance =  $\pm 20\%$ Heavy recoil tagging in fusion-evaporation and transfer reaction spectroscopy Applications: nuclear astrophysics and exotic nuclear structure Commissioning in 2011

#### Probing the Transition to the Island of Inversion TIGRESS & SHARC (C. Diget et al.)



At ISAC2 we will use the <u>isotone</u> of <sup>24</sup>Ne namely <sup>25</sup>Na as the projectile.

We aim to <u>test the modifications to USD</u> that have been used to reproduce the raised d3/2 level seen in <sup>25</sup>Ne, by measuring the <u>p-n coupling states</u> in <sup>26</sup>Na.

The experiment will use 10<sup>6</sup> pps <sup>25</sup>Na at 5.00 MeV/u.





# The future site map of TRIUMF



#### Proposal:

•BL4N is proposed to deliver 500-MeV protons to two actinide target stations for beam production

•Take advantage of the shielded and unused proton hall to add a 50-MeV electron driver to supply electrons to the new target area via a separate beamline

•Develop new ISAC front end to permit three simultaneous RIB beams (two accelerated)

•E-linac CFI funded (awaiting matching funds)



## The People (some)

Alan Astbury, Alan Shotter and Erich Vogt Paul Schmor Marik Dombsky – ISOL Targetry Pierre Bricault- ISOL ion sources and other components Lothar Buchmann (w Dick Azuma, Jim King) – Reactions with RB for Astrophysics Jack Beveridge – Shielding and initial layout Harvey Schneider (and Bob Laxdal and PB) – RFQ and LINAC Otto Hauser (and Peter Jackson and John Behr) – TRINAT Dave Hutcheon – DRAGON (Technical aspects) Gordon Ball – Nuclear Structure and, and, and Rob Kiefl (Phil Levy,Gerald Morris)–Condensed Matter (with polarized <sup>8</sup>Li) Jean Michel Poutissou – Science leadership and beam scheduling

Newer group Jens Lassen Jens Dilling Chris Ruiz Barry Davids Carl Svenson (and Greg Hackman and Paul Garrett) Matt Pearson

Entire TRIUMF staff (Engineers, Remote Handling Group, Technicians, Machinists, Draftspersons and Designers, etc...)

# **Concluding Remarks**

- ISAC is now a world class facility for nuclear astrophysics, nuclear structure studies, fundamental symmetries and condensed matter physics;
- Window of opportunity before FRIB, FAIR and others;
- Needs constant development of new radioactive and stable heavy ion beams and with higher intensities;
- First class training ground for students and post-docs;
- Bright future with existing world class, unique facilities;
- Exciting forward looking 5 year plan;
- BUT, should not lose sight of what got it here:

- an excellent scientific, technical and professional staff (dedicated, open, cooperative, loyal, good will, creative)

# End of Slides

## FRIB at MSU



## **SPIRAL 2**





## E-linac: MW-class Superconducting Electron Accelerator at TRIUMF



- A MW-class electron linac is a driver for photofission with rates up to 10<sup>14</sup> fissions/sec.
- The present e-linac design concept, based on 1.3 GHz SRF CW operation, offers flexibility, possibility for expansion to other applications (Free Electron Laser, Energy Recovery Linac).



Design parameters: 50 MeV, 10 mA, CW operation.

## $\beta$ -NMR probes phase changes





- Observed solar abundances of Cd/In/Pd, Hf/Ta/W not compatible with nuclear shell model!
- Quenching shell gaps in calculations gives answer closer to abundances
   21Na Matrix R

21Na Matrix Elements: First TIGRESS RIB





- Isotope shift measurements: ToPLiS collaboration @ ISAC measured laser frequency shifts for the lithium isotopes
- G. W. Drake (Windsor) PRL. 100, 243002 (2008) atomic theory calculations for the mass shifts => extract the charge radius
- Isotope shift = modification of electron binding energy =Mass Shift (mass effect) + Field shift (finite size of nucleus)






## ISAC: The First 10 Years

- What is ISAC
- What is its place in the World
- How did it start
  - Proposal in 1985
  - TISOL (10 year for learning)
    - Key Studies at TISOL
  - Buiding the facility
    - Remote Handling and Shielding
    - The Target and Beam Production
    - Key Components
      - The RFQ Accelerator
      - The High Resolution Mass Anayzer
    - ISAC II
- Science in the First 10 Years
  - Nuclear Astrophysics
    - DRAGON
    - TUDA
    - Other studies (Ne)
  - Fundamental Physics
    - TRINAT
    - Superallowed Beta Decay
  - Condensed Matter Physics
    - Beta NMR and Polarized <sup>8</sup>Li
  - Nuclear Physics and Structure
    - The 8 Pi
    - TITAN
    - TIGRESS
    - Li11 Charge Distribution
- What does the future hold?
- Concluding remarks (people, creativity, good will and cooperation)

# Planned Experimental Facilities and Beams

- EMMA
- TIGRESS (Completed)
- Actinide Targets (on a regular basis) TRINAT and FrPNC Radon EDM r-process masses for astrophysics
- Second Production Target System including e-LINAC



## The TRIUMF-ISAC Radioactive Beams Facility

- ISAC I Project proposed in 1985; funded in 1995 (5 year plan)
- RB Production by the ISOL Method (500 MeV p<sup>+</sup>)
- RB Accelerated using LINACS (0.15 1.5 MeV/u): ISAC I
- Two ISAC 1 Experimental Areas (LEBT and HEBT)
- ISAC II funded in 2000;
- Major Technical Milestones
  - 1998 First RB beam (<sup>38m</sup>K) to TRINAT
  - 2000 First physics (<sup>74</sup>Rb lifetime with high precision)
  - 2001 TUDA and DRAGON perform RB experiments;<sup>21</sup>Na
  - 2002  $8\pi$  and  $\beta$ -NMR perform physics
  - 2003 TITAN and TIGRESS Funded; ISAC II bldg. opened
  - 2004 ECR used for exp.; CSB tested; <sup>26</sup>Al beam;
    - (high power target, <sup>11</sup>Li; laser ion source)
  - $-200X {}^{26g}Al exp.$ ,  ${}^{11}Li CSD done$ , + many other exps.

### Radioactive Beams at TRIUMF-ISAC The ISOL Method

- 500 MeV protons onto thick target
- Spallation, fragmentation, (fission) reactions
- Have used Nb, Ta, SiC, TiC, CaO, CaZrO<sub>3</sub>, (ZrC)
- Intensities up to 150  $\mu$ A possible (now 100  $\mu$ A)
- Products diffuse out at high temperatures
- Species ionized in heated surface and laser ion sources; ECR (2004, revised 2008); FEBIAD(2006)
- Beams delivered to ISAC II (2006-7)
- Successfully tested uranium (actinide) target 2008

#### **Some Beam Intensities at Yield Station**

<sup>8</sup> Li	(Ta)	8 x 10 <sup>8</sup> pps
<sup>11</sup> Li	(Ta)	4 x 10 <sup>4</sup> pps
<sup>21</sup> Na	(SiC)	9.9 x 10 <sup>9</sup> pps
<sup>26g</sup> Al	(SiC)	1 x 10 <sup>10</sup> pps
<sup>74</sup> Rb	(Nb)	1.3 x 10 <sup>4</sup> pps
<sup>79</sup> Rb	(Nb)	4.6 x 10 <sup>9</sup> pps
<sup>160</sup> Yb	(Ta)	8.4 x 10 <sup>9</sup> pps

M. Dombsky TRIUMF www.triumf.ca/people/marik/