

# Mass measurement of neutron-rich isotopes around $N = 32$ and $N = 34$

K, Ca, Sc RIB & TITAN

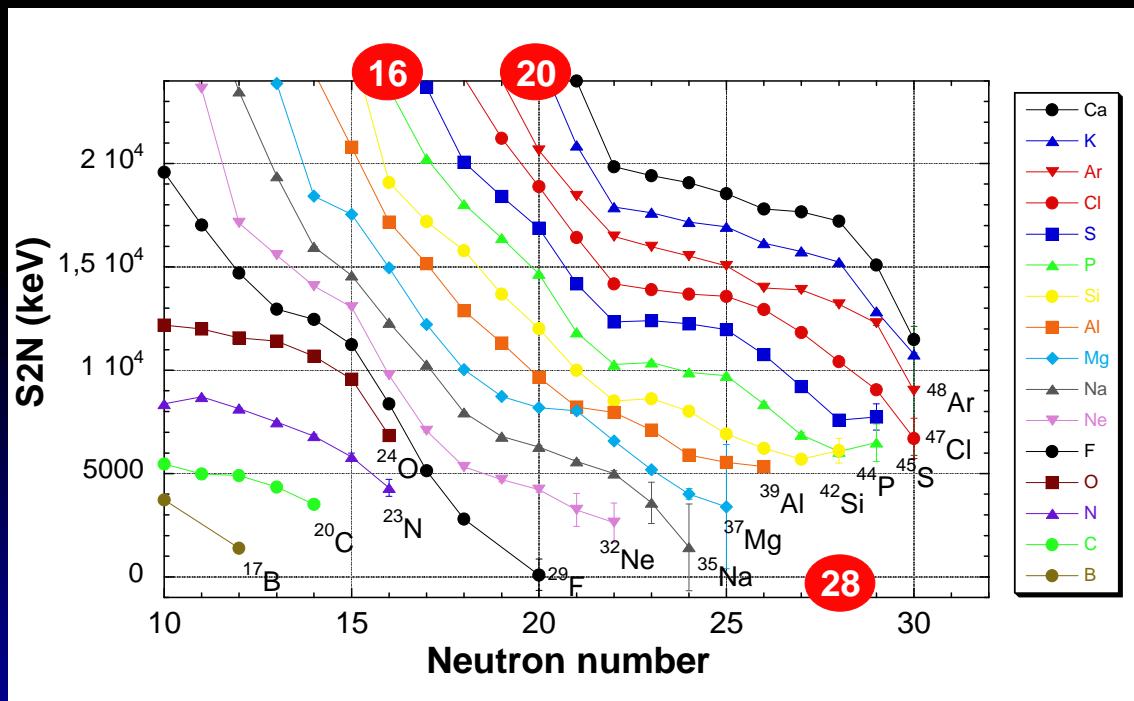
*H. Savajols (GANIL/TRIUMF) & J. Dilling (TRIUMF)*

# Nuclear structure in neutron-rich nuclei

## Mass measurement

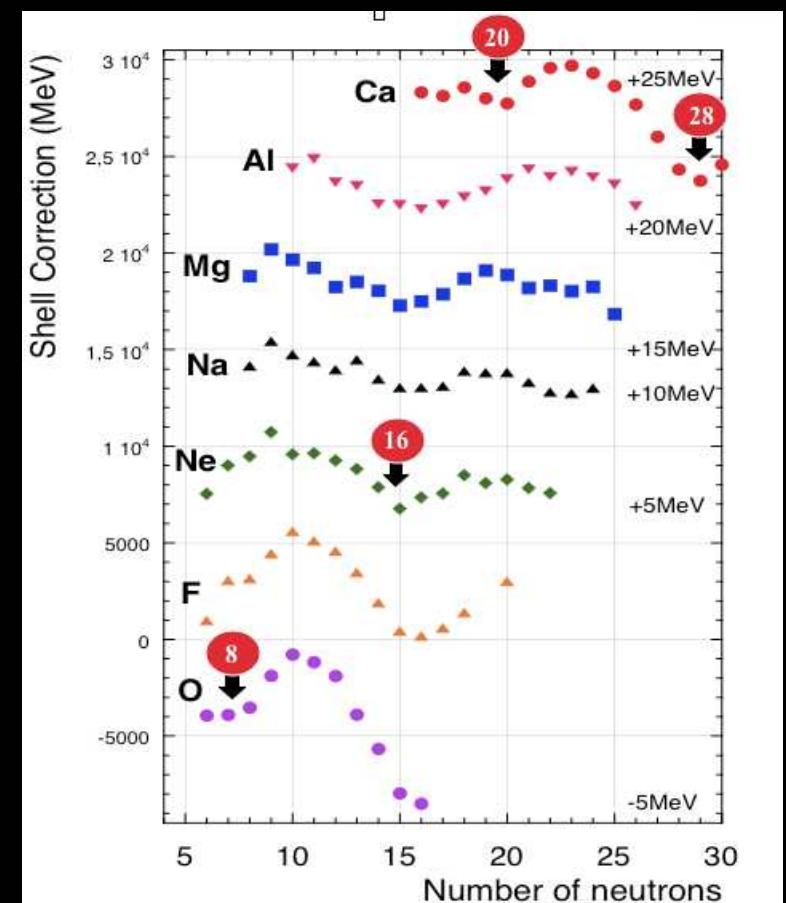
⇒ a powerful nuclear structure probe

$$S_{2n}(A,Z) = M(A-2,Z) - M(A,Z) + 2M_n$$



$$\text{Shell Correction} = \Delta M_{\text{exp}} - \Delta M_{\text{FRLLDM}}$$

P.Moller, J.R.Nix et al., Atomic Data and Nuclear Data Tables 59 (1995) 185

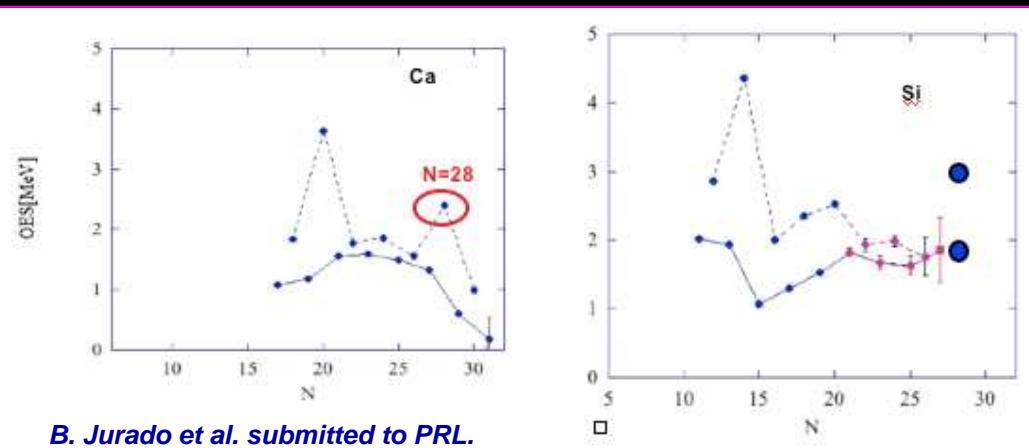


# Mass measurement $\rightarrow$ a powerful nuclear structure probe

$$\Delta_3(N) = (-1)^N [M(N-1) + M(N+1) - 2M(N)] k^2/2$$

- Pairing effects  $\downarrow \Delta_3$  at odd values of N
- Single particle spacing  $\downarrow$  difference of  $\Delta_3$  at adjacent even and odd values of N

*Satula et al., PRL 81 (1998) 3599*

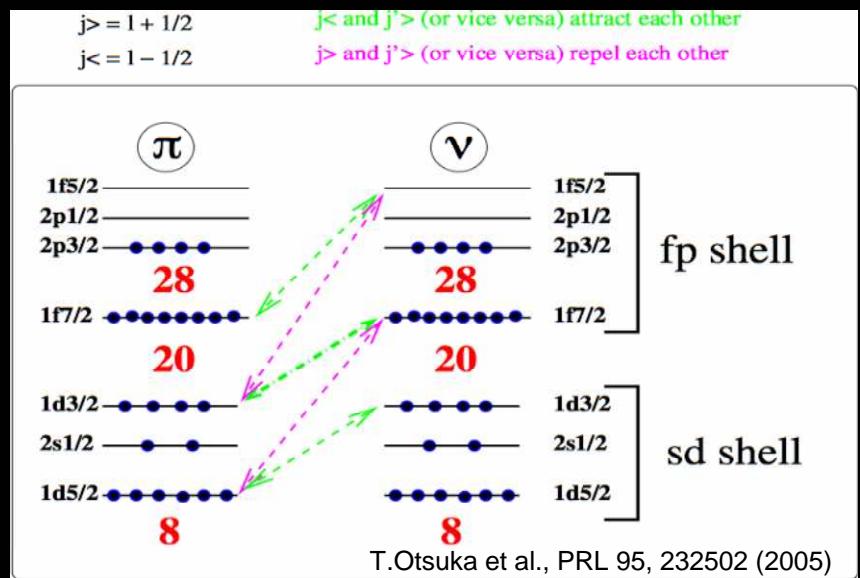


*B. Jurado et al. submitted to PRL.*

## ⇒ On neutron rich-side

▀ N=16, N=20 and N=28

Clear evidences on the change in the shell structure

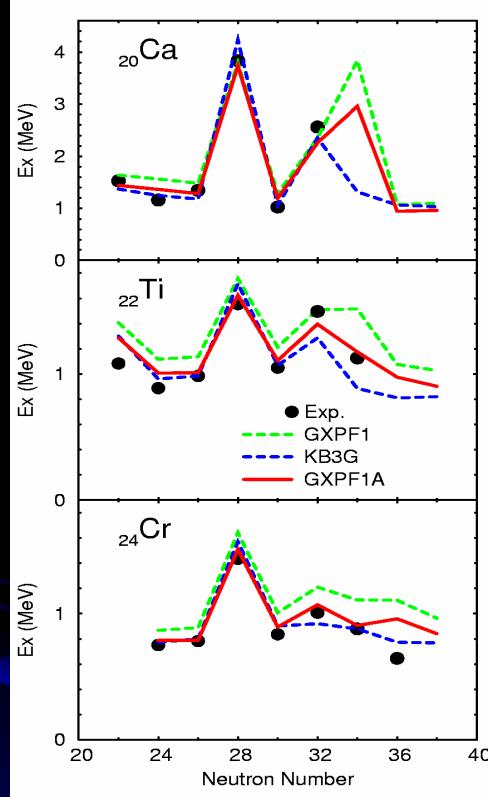


Manifestation of the tensor monopole part of the NN interaction in  $N/Z >> 1$  Nuclei ?

## ⇒ Inversion of shell

Dependence of the  $\nu-\pi$  interaction of their combination of their spin

# E1112 : mass measurement around N=32 and N=34



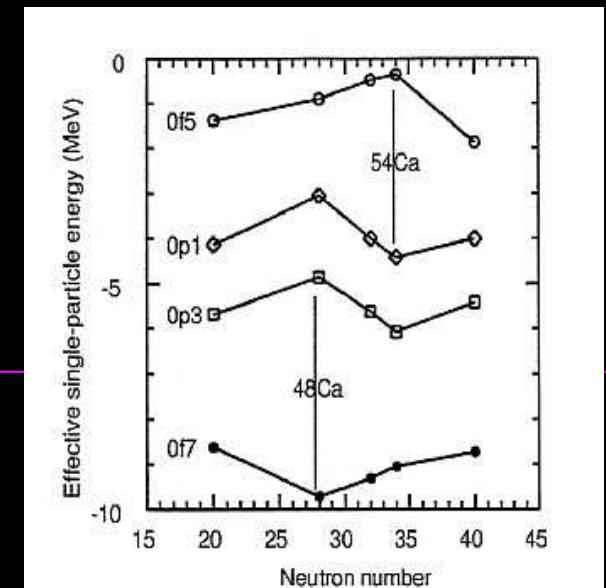
## In the *pf* shell :

- The relative energy of the  $p_{3/2}$ ,  $p_{1/2}$  and  $f_{5/2}$  orbits determine where sub shell closure take place
- Most of the shell model effective interaction predict a sub shell at N=32 for the Ca isotope

(confirm by  $\beta$ -decay measurement,  $B(E2)$ , high spin of the even-even  $^{56}\text{Cr}$  and  $^{54}\text{Ti}$ )

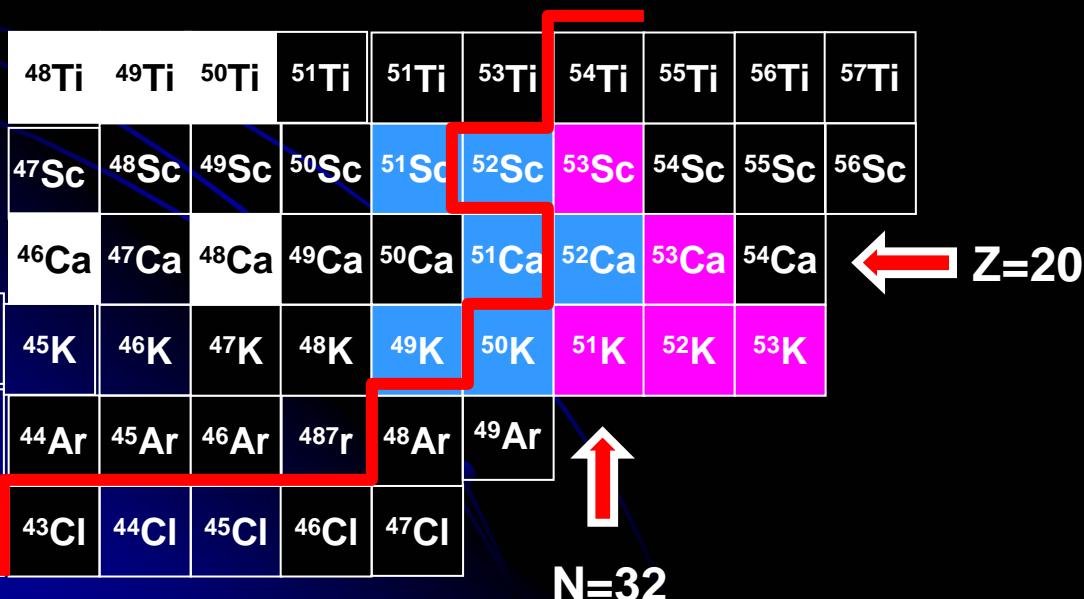
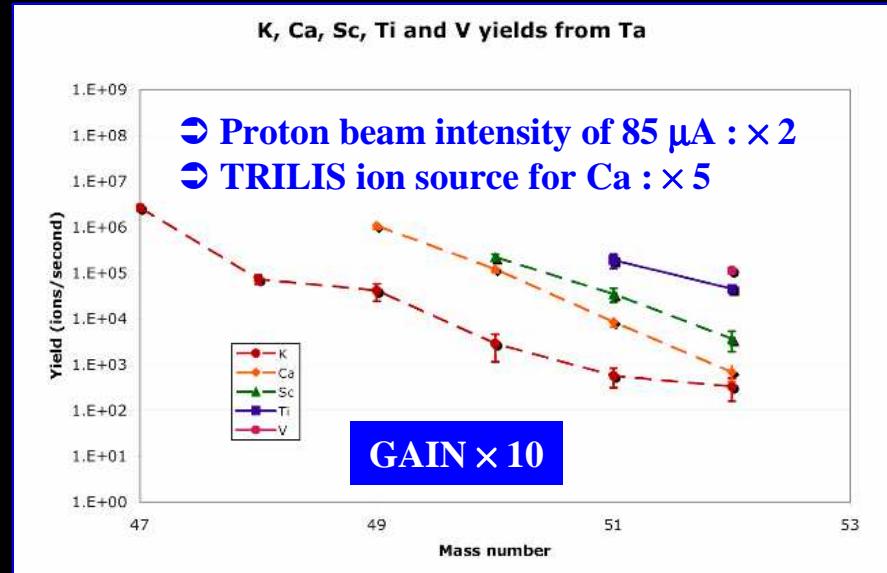
- New effective interaction :  
GXPF1 Homma et al. PRC65, (2002) 1301R  
Sizable energy gap between  $p_{1/2}$  and  $f_{5/2}$  orbits  $\downarrow N=34$   
neutron proton interaction  $\pi f_{7/2} - \nu f_{5/2}$

- Shell-structure evolution is expected to be reflected in mass measurements



# Proposed experiment :

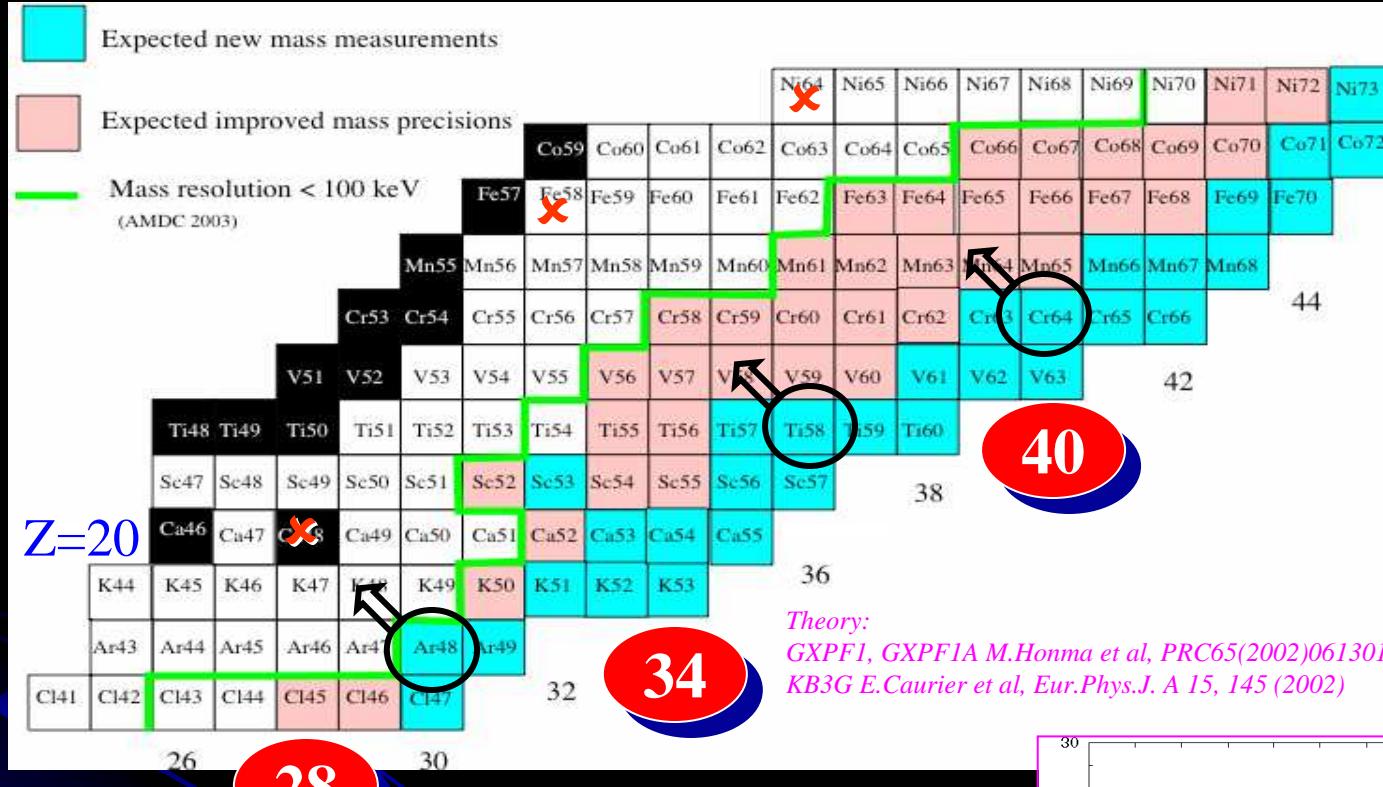
| <i>Isotope</i>   | <i>Half-life</i> | <i>(Expected) Yield</i> | <i>Ion source</i> |
|------------------|------------------|-------------------------|-------------------|
| $^{49}\text{K}$  | 1.26 s           | $2 \times 10^6$         | Surface           |
| $^{50}\text{K}$  | 472 ms           | $1 \times 10^6$         | Surface           |
| $^{51}\text{K}$  | 365 ms           | $2 \times 10^6$         | Surface           |
| $^{52}\text{K}$  | 105 ms           | $1 \times 10^6$         | Surface           |
| $^{53}\text{K}$  | 30 ms            | $5 \times 10^6$         | Surface           |
| $^{51}\text{Ca}$ | 10 s             | $9 \times 10^6$         | TRILIS            |
| $^{52}\text{Ca}$ | 4.6 s            | $8 \times 10^6$         | TRILIS            |
| $^{53}\text{Ca}$ | 90 ms            | $7 \times 10^6$         | TRILIS            |
| $^{51}\text{Sc}$ | 12.4 s           | $1 \times 10^6$         | Surface           |
| $^{52}\text{Sc}$ | 8.6 s            | $8 \times 10^6$         | Surface           |
| $^{53}\text{Sc}$ | > 3 s            | $1 \times 10^6$         | Surface           |



- Expected improved mass precisions
- Expected new mass precisions
- Mass resolution < 100 keV (AMDC 2003)

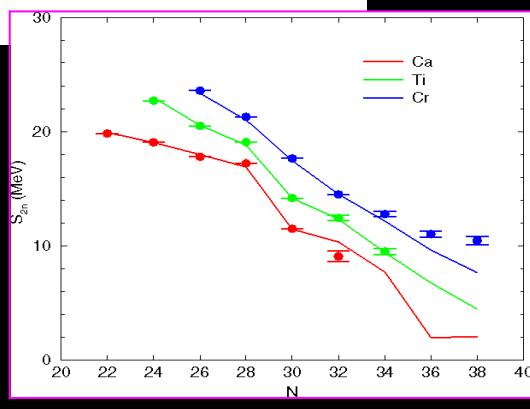
11 new masses of very exotic nuclei  
○ 5 new masses  
○ 6 improved masses

# E418a : measurements of 31 new masses and 37 improved masses

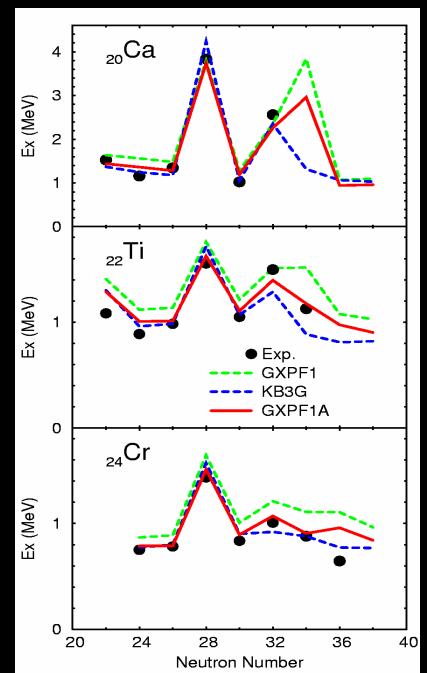


## Shell structure

- beyond N=28
- Around N=40
- New magic number N=34 ?  
(neutron proton interaction  $\pi f_{7/2} - vf_{5/2}$ )



T.Otsuka PRL 87 (2001) 082502-1  
And private communication



## Astrophysical context

■  $S_n$  important to determine the r-process path



turning points ( $t_\beta$  vs  $t_n$ )

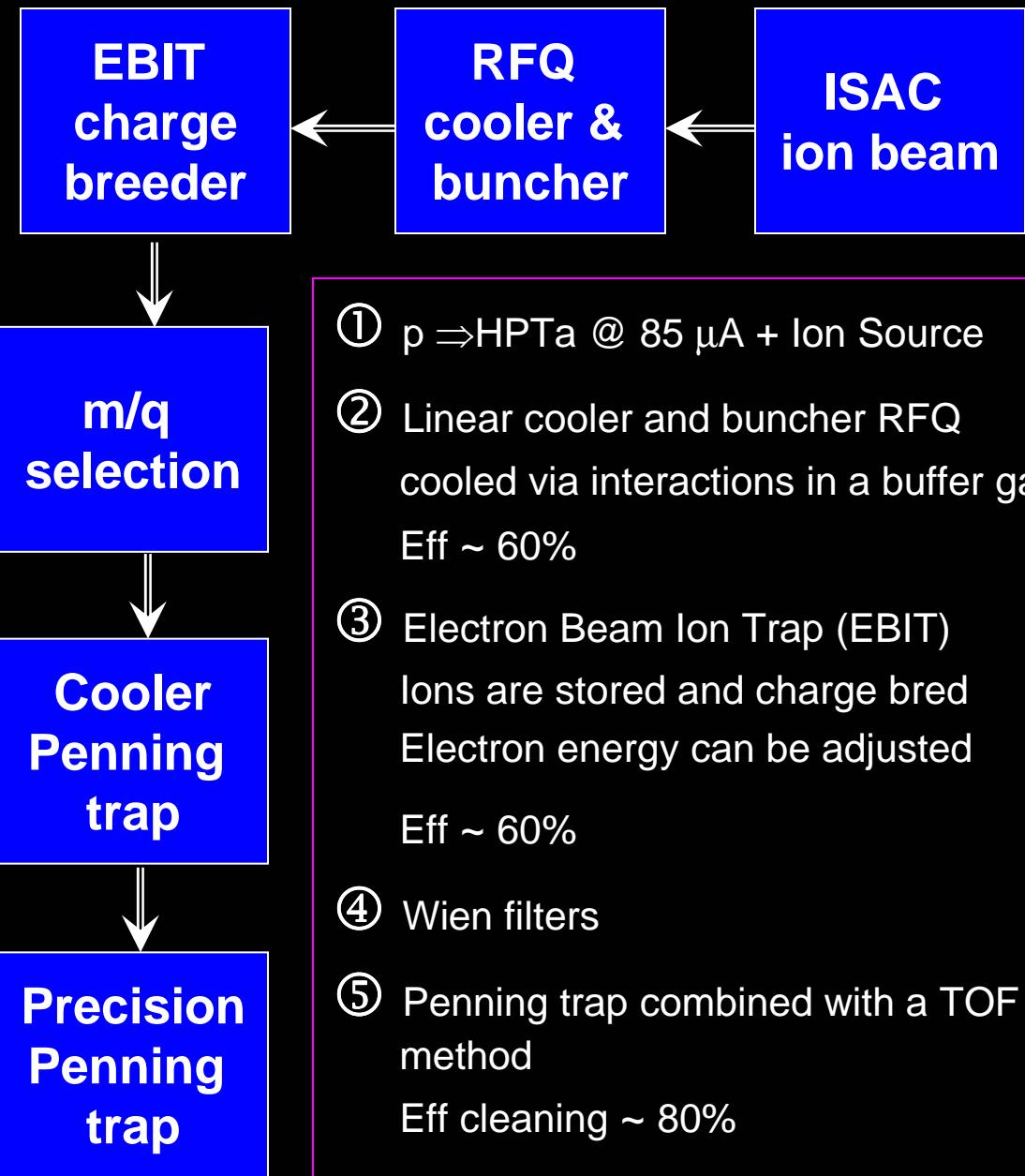
O.Sorlin et al., NPA A660 (1999) 3.  
S.Grevy, private communication

# TITAN device

Mass measurements on isotopes with short half-life  $T_{1/2}$  10 ms and low production yields ( $\approx 100$  ions/s) with high precision  $\delta m/m \approx 10^{-9}$ .

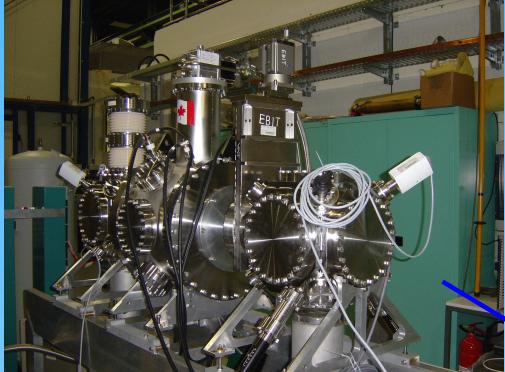
Ideally, uniquely matched to isotope type production mode.

TITAN started April 2003, planned first on-line mass measurements will be in 2006.



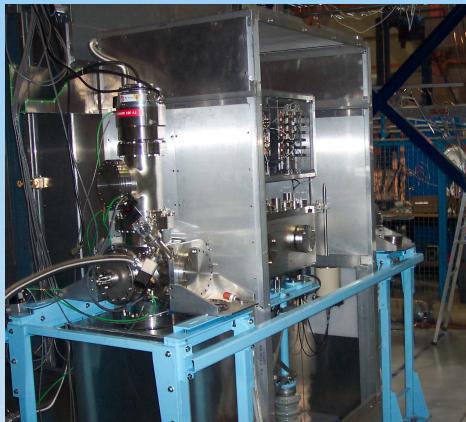


Max-Planck-Institut  
für Kernphysik



EBIT built @ MPI-HD.  
to TRIUMF April 2006.

Cooler trap for HCI  
(to be built in Manitoba,  
CFI grant received)

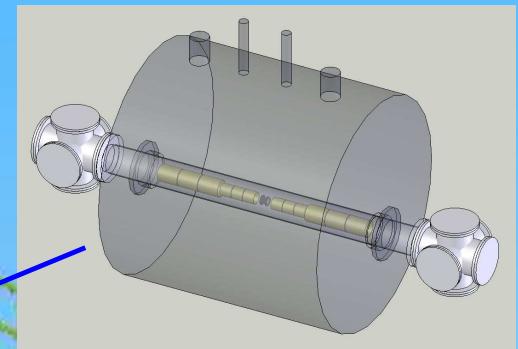
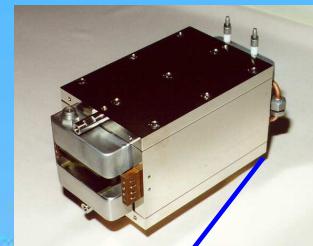


RFQ operational on test bench  
Moved to ISAC, ready for tests

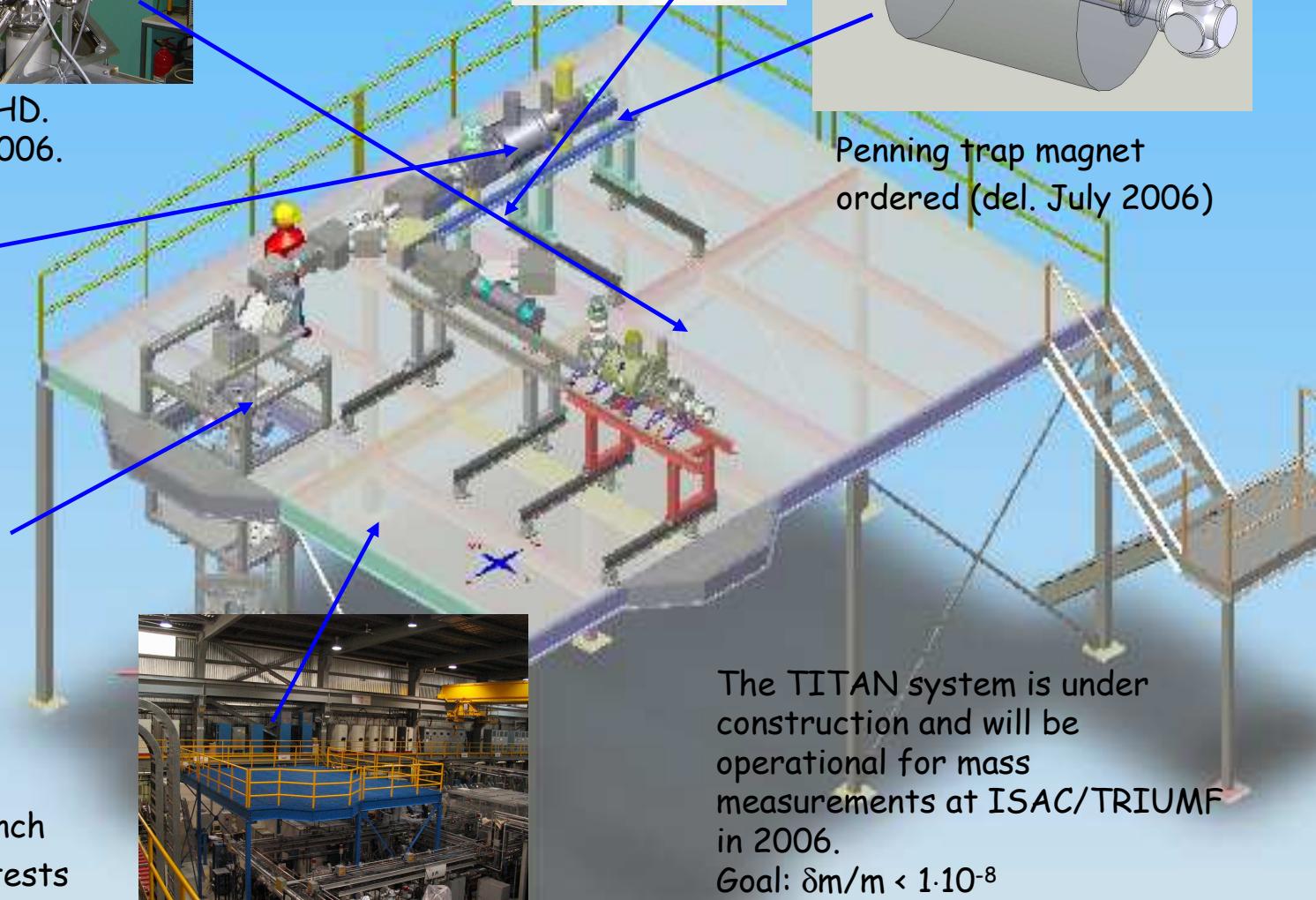
TRIUMF



Wien filter  
(R=500)



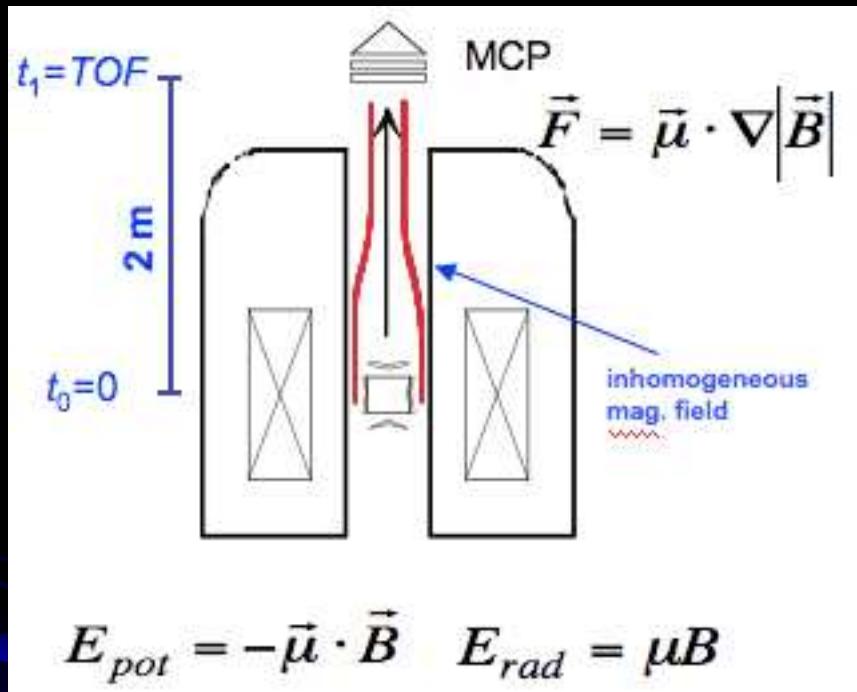
Penning trap magnet  
ordered (del. July 2006)



TITAN platform finished at ISAC

The TITAN system is under  
construction and will be  
operational for mass  
measurements at ISAC/TRIUMF  
in 2006.  
Goal:  $\delta m/m < 1 \cdot 10^{-8}$

# Mass measurement via time-of-flight



QuickTime™ and a  
TIFF (LZW) decompressor  
are needed to see this picture.

Determine atom mass from frequency ratio with a well known reference

# Beam time request

| Ions             | Half-life | Intensity | Breeding to He like | e <sup>-</sup> energy | Measuring | Resolution         |
|------------------|-----------|-----------|---------------------|-----------------------|-----------|--------------------|
| <sup>53</sup> Sc | 0.9 s     | 700/s     | Q = 19 - dt = 30 ms | 4500 eV               | 100 ms    | 5.10 <sup>-9</sup> |
| <sup>53</sup> Ca | 90 ms     | 700/s     | Q = 18 - dt = 30 ms | 4100 eV               | 100 ms    | 5.10 <sup>-9</sup> |
| <sup>53</sup> K  | 30 ms     | 500/s     | Q = 17 - dt = 30 ms | 3600 eV               | 30 ms     | 1.10 <sup>-8</sup> |

|                  |                  |                  |                  |                  |                  |                  |                  |                  |                  |
|------------------|------------------|------------------|------------------|------------------|------------------|------------------|------------------|------------------|------------------|
| <sup>48</sup> Ti | <sup>49</sup> Ti | <sup>50</sup> Ti | <sup>51</sup> Ti | <sup>51</sup> Ti | <sup>53</sup> Ti | <sup>54</sup> Ti | <sup>55</sup> Ti | <sup>56</sup> Ti | <sup>57</sup> Ti |
| <sup>47</sup> Sc | <sup>48</sup> Sc | <sup>49</sup> Sc | <sup>50</sup> Sc | <sup>51</sup> Sc | <sup>52</sup> Sc | <sup>53</sup> Sc | <sup>54</sup> Sc | <sup>55</sup> Sc | <sup>56</sup> Sc |
| <sup>46</sup> Ca | <sup>47</sup> Ca | <sup>48</sup> Ca | <sup>49</sup> Ca | <sup>50</sup> Ca | <sup>51</sup> Ca | <sup>52</sup> Ca | <sup>53</sup> Ca | <sup>54</sup> Ca |                  |
| <sup>44</sup> K  | <sup>45</sup> K  | <sup>46</sup> K  | <sup>47</sup> K  | <sup>48</sup> K  | <sup>49</sup> K  | <sup>50</sup> K  | <sup>51</sup> K  | <sup>52</sup> K  | <sup>53</sup> K  |
| <sup>43</sup> Ar | <sup>44</sup> Ar | <sup>45</sup> Ar | <sup>46</sup> Ar | <sup>48</sup> Ar | <sup>48</sup> Ar | <sup>49</sup> Ar |                  |                  |                  |
| <sup>42</sup> Cl | <sup>43</sup> Cl | <sup>44</sup> Cl | <sup>45</sup> Cl | <sup>46</sup> Cl | <sup>47</sup> Cl |                  |                  |                  |                  |

Summarizing the duty cycle in the case of <sup>53</sup>Sc and for 700 incoming ions /s,

- 100 ms cooling × 60% efficiency × decay losses = 36 ions
- 30 ms breeding × 60% efficiency × decay losses = 21 ions
- 100 ms cleaning × 80% efficiency × decay losses = 14 ions
- 100 ms measuring × 50% efficiency × decay losses = 6 ions

For <sup>53</sup>Ca & <sup>53</sup>K, the cycle has to be optimised to get 1 ion/cycle

1 ion/cycle → 3 ions/s

N = 3000

17 mn for one measurement !

- 3 shifts for the set-up of each elements
- 1 shift for each isotope

Total of 20 shifts

# Collaboration

|                   |                          |                    |      |
|-------------------|--------------------------|--------------------|------|
| H. Savajols       | GANIL/TRIUMF             | Scientist          | 50 % |
| J. Dilling        | TRIUMF                   | Scientist          | 50 % |
| P. Delheij        | TRIUMF                   | Scientist          | 50 % |
| W. Mittig         | GANIL                    | Scientist          | 20 % |
| P. Roussel-Chomaz | GANIL                    | Scientist          | 20 % |
| A.C.C. Villari    | GANIL                    | Scientist          | 20 % |
| A. Chbihi         | GANIL                    | Scientist          | 10 % |
| A. Gillibert      | CEA/Saclay               | Scientist          | 10 % |
| N. Orr            | LPC/Caen                 | Scientist          | 10 % |
| M. Chartier       | Univ. Liverpool          | Scientist          | 10 % |
| G. Ball           | TRIUMF                   | Scientist          | 10 % |
| G. Hackman        | TRIUMF                   | Scientist          | 10 % |
| I. Tanihata       | TRIUMF                   | Scientist          | 10 % |
| C. Ravuri         | TRIUMF/SFU               | Research Associate | 10 % |
| R. Kanungo        | TRIUMF                   | Research Associate | 10 % |
| F. Sarazin        | Colorado School of Mines | Professor          | 10 % |
| G. Gwinner        | U. of Manitoba           | Professor          | 10 % |
| A. Allaoua        | GANIL                    | Phd Student        | 50 % |
| L. Gaudefroy      | GANIL                    | Post-Doc           | 10 % |
| M. Smith          | UBC                      | Student            | 50 % |
| M. Brodeur        | UBC                      | Student            | 50 % |
| C. Champagne      | McGill                   | Student            | 50 % |
| V. Ryjkov         | TRIUMF                   | Post-Doc           | 50 % |