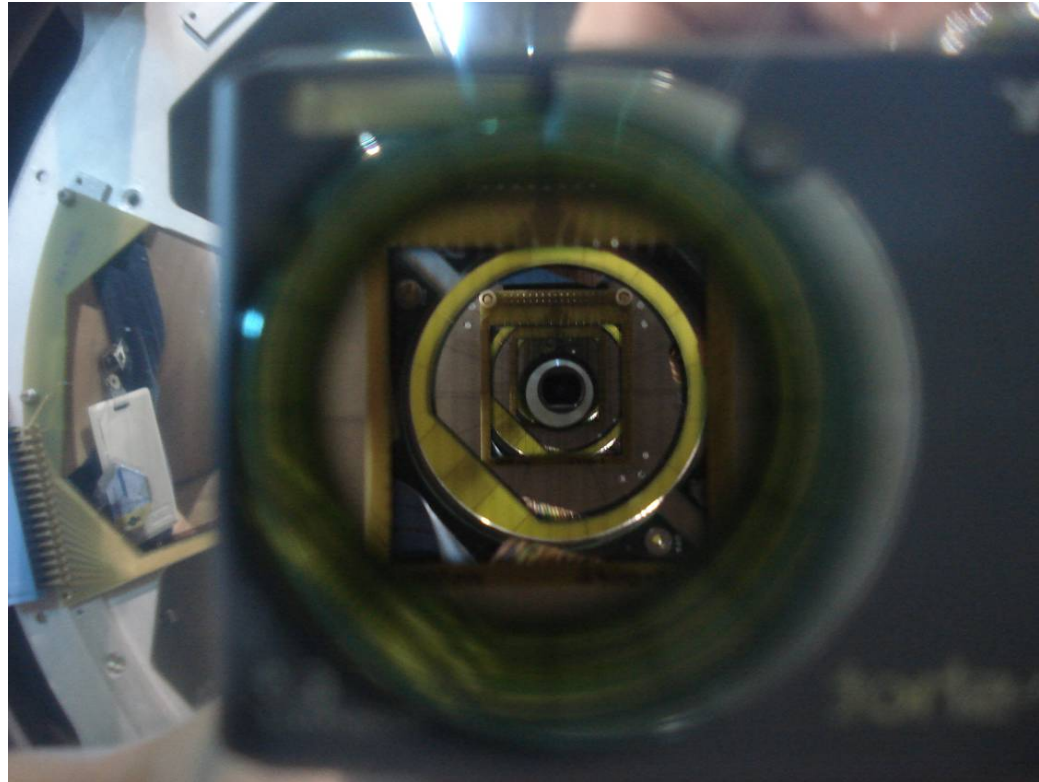


S1103 - (α ,p) reactions in type I X-ray bursts:
time-reversed approach at ISAC II



Marialuisa Aliotta

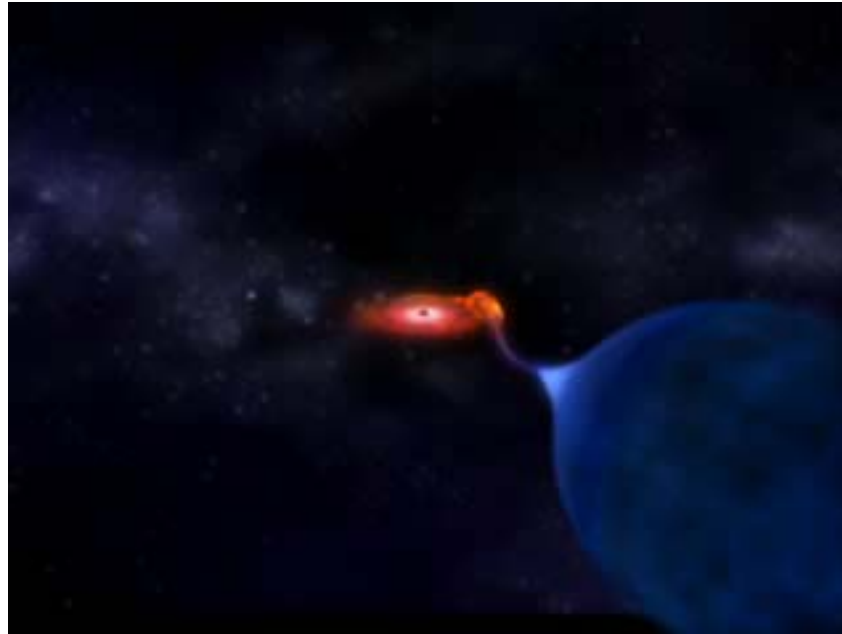


School of Physics and Astronomy - University of Edinburgh
Scottish Universities Physics Alliance



X-Ray Bursts

semi-detached binary system: Neutron star + less evolved star



- thermonuclear runaway on surface of accreting neutron star
- explosive ignition of H and/or He
- emission of strong X-rays
- if matter transfer continues process may repeat

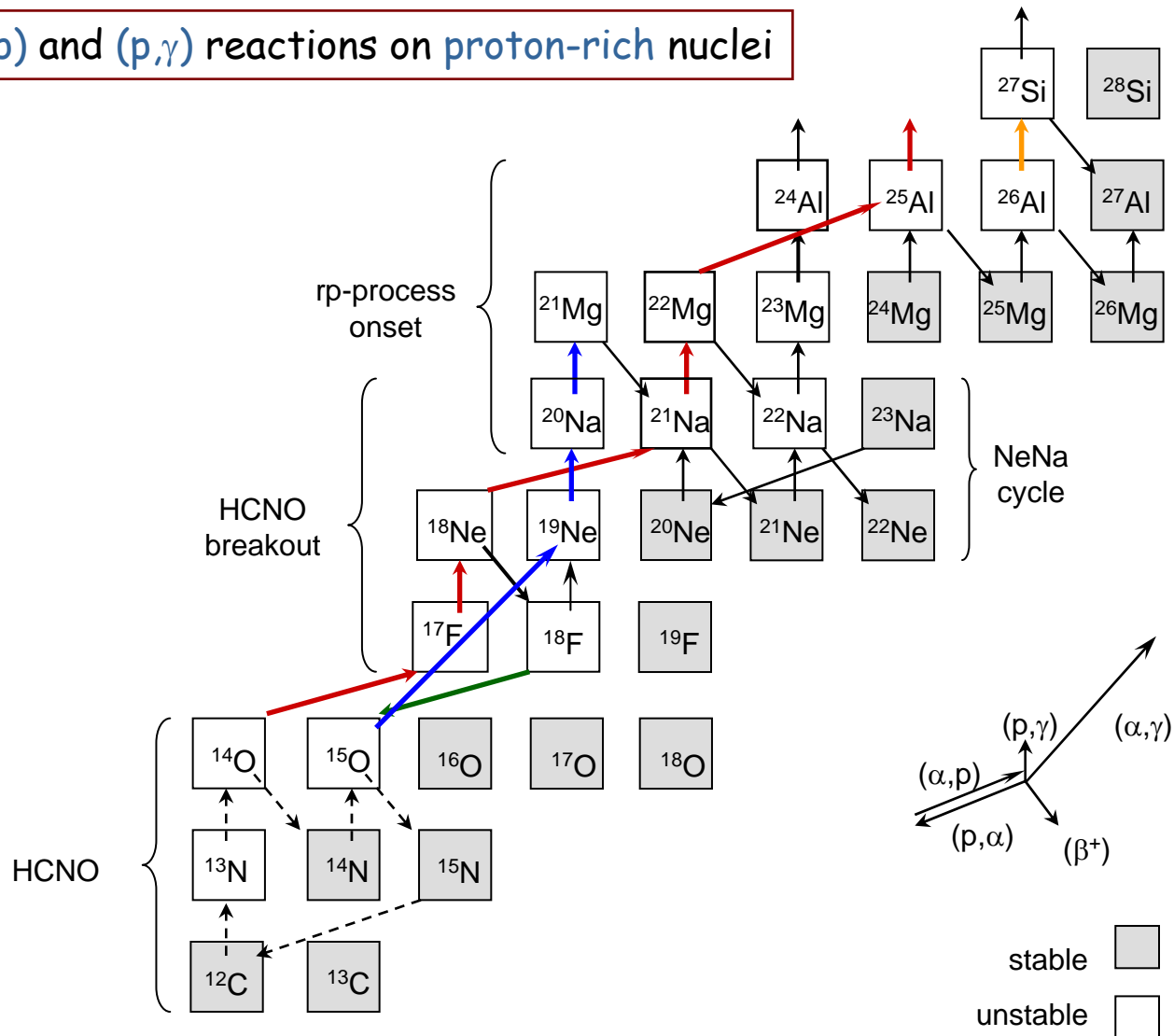
explosive hydrogen burning

$T \sim 10^9 \text{ K}$
 $\rho \sim 10^6 \text{ g cm}^{-3}$



nucleosynthesis up to $A \sim 80-100$ mass region

(α, p) and (p, γ) reactions on proton-rich nuclei



(exact path depends on given stellar conditions)

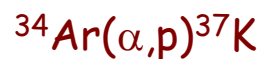
proposal

study key (α ,p) reactions relevant to type I X-ray bursts



breakout from HCNO cycle

direct and indirect investigations exist, but uncertainties remain



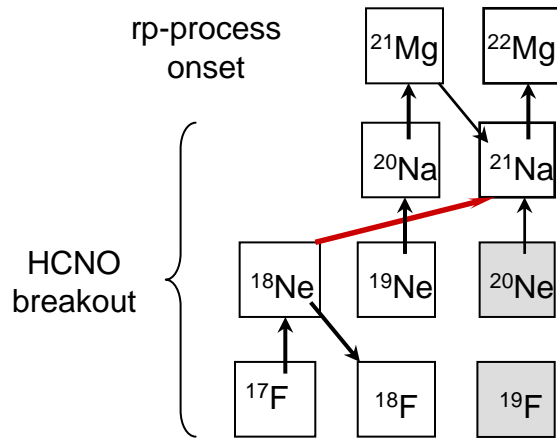
} possible waiting points in type I X-ray bursts

no experimental data available

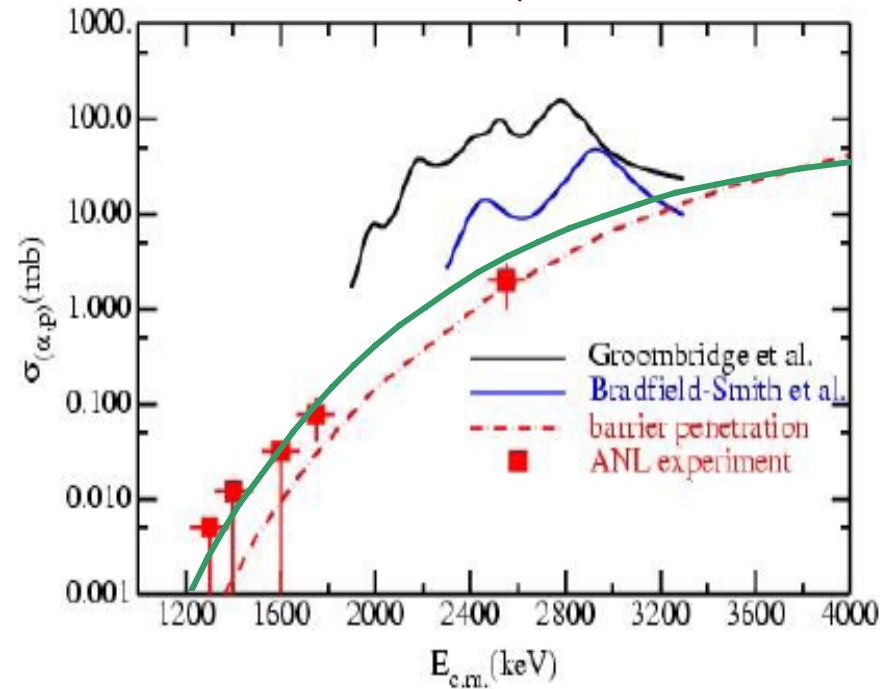
- direct studies hampered both by low RIB intensities and low target densities
- propose to study *time-reversed* (p,α) processes at relevant astrophysical energies

$^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$

one of two main **breakout routes**
from HCNO cycle into **rp-process**



Argonne National Laboratory
Internal report 2005



current status:

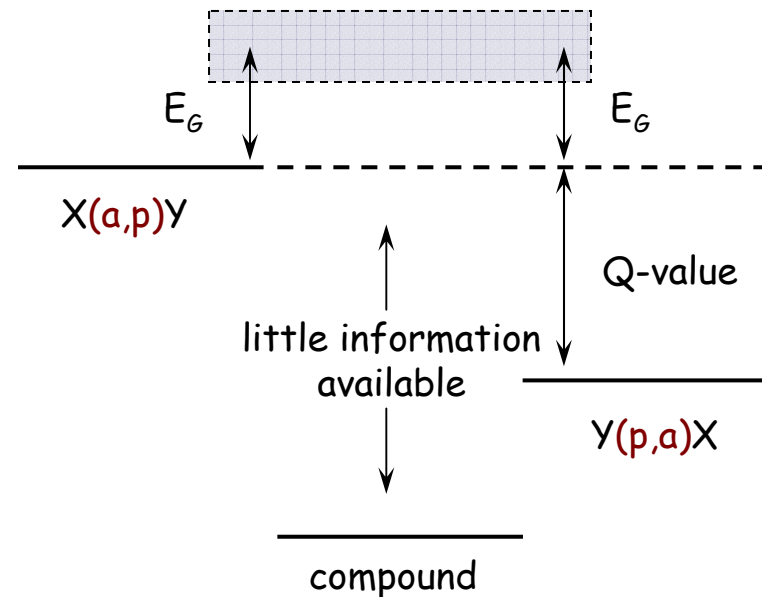
- two **direct** measurements
- one **time-reversed** measurement
- theoretical calculation (**Hauser-Feshbach**)
- large **discrepancies**

the methodology

time-reversed approach: $X(\alpha,p)Y \Leftrightarrow Y(p,\alpha)X$

detailed-balance theorem

$$\frac{\sigma_{\alpha X}}{\sigma_{pY}} = \frac{m_p m_Y}{m_\alpha m_X} \frac{E_{pY}}{E_{\alpha X}} \frac{(2J_p + 1)(2J_Y + 1)}{(2J_\alpha + 1)(2J_X + 1)}$$



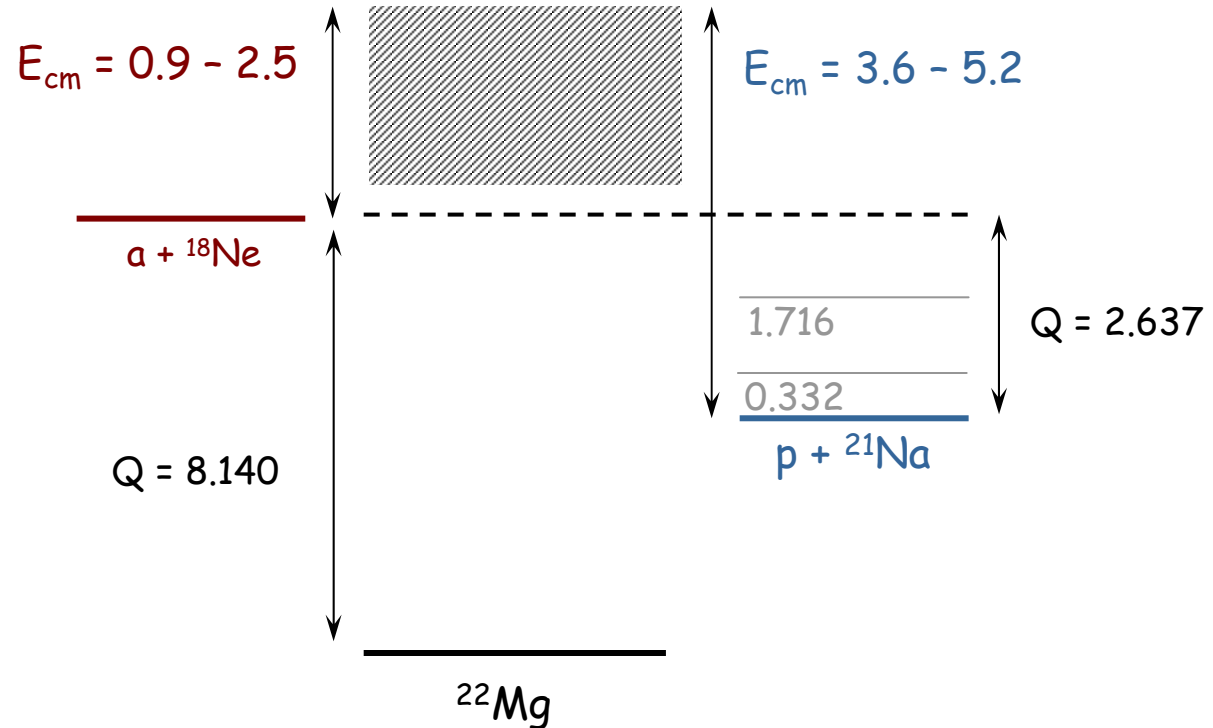
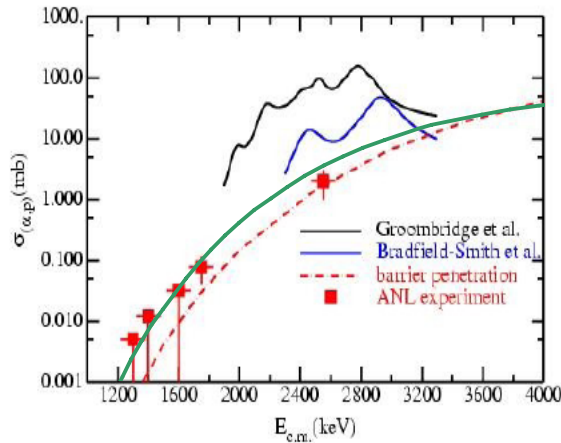
direct approach: spin-less particles \Rightarrow only **natural parity** states can be populated

indirect approach: no selectivity

However! kinematic selection on transitions between ground states
 \Rightarrow ensure only natural parity states have been populated

the case for $^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$

(all energies in MeV)



- Aims:
- investigate energy range $E_{cm} \sim 0.9 - 2.5$ MeV by time-reverse approach ($E_{cm} \sim 3.6 - 5.2$ MeV)
 - investigate resonant states in ^{22}Mg ($E_r \sim 8.5 - 10.1$ MeV)

Experimental setup @ ISAC II

$^{21}\text{Na } 5^+$ beam
 $I \sim 5 \times 10^6$ pps
 $E \sim 5.5 - 3.8$ MeV/u

W: protons (alphas)

elastic proton scattering
 measurement

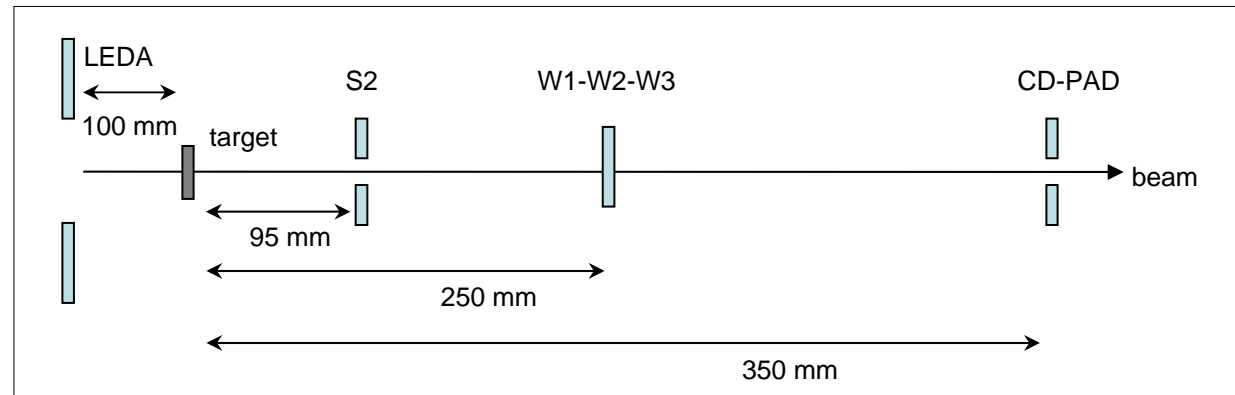
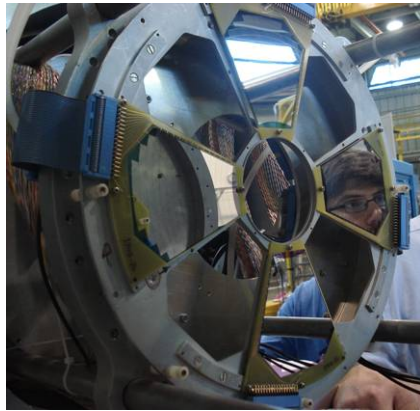
CH2 target



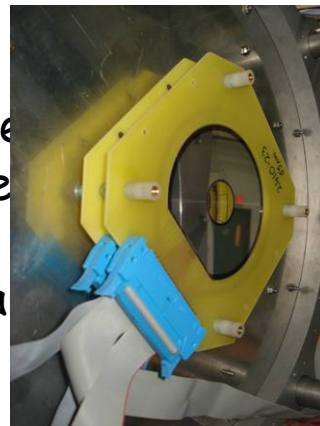
aim 2
 meas
 (thick)



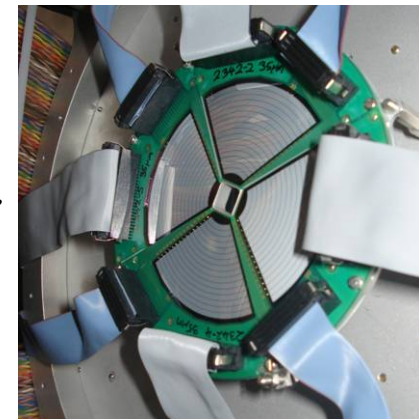
LEDA: RBS on Au spot

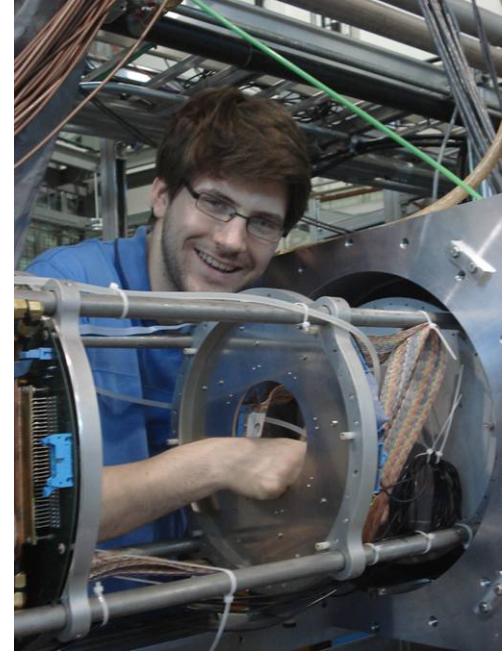
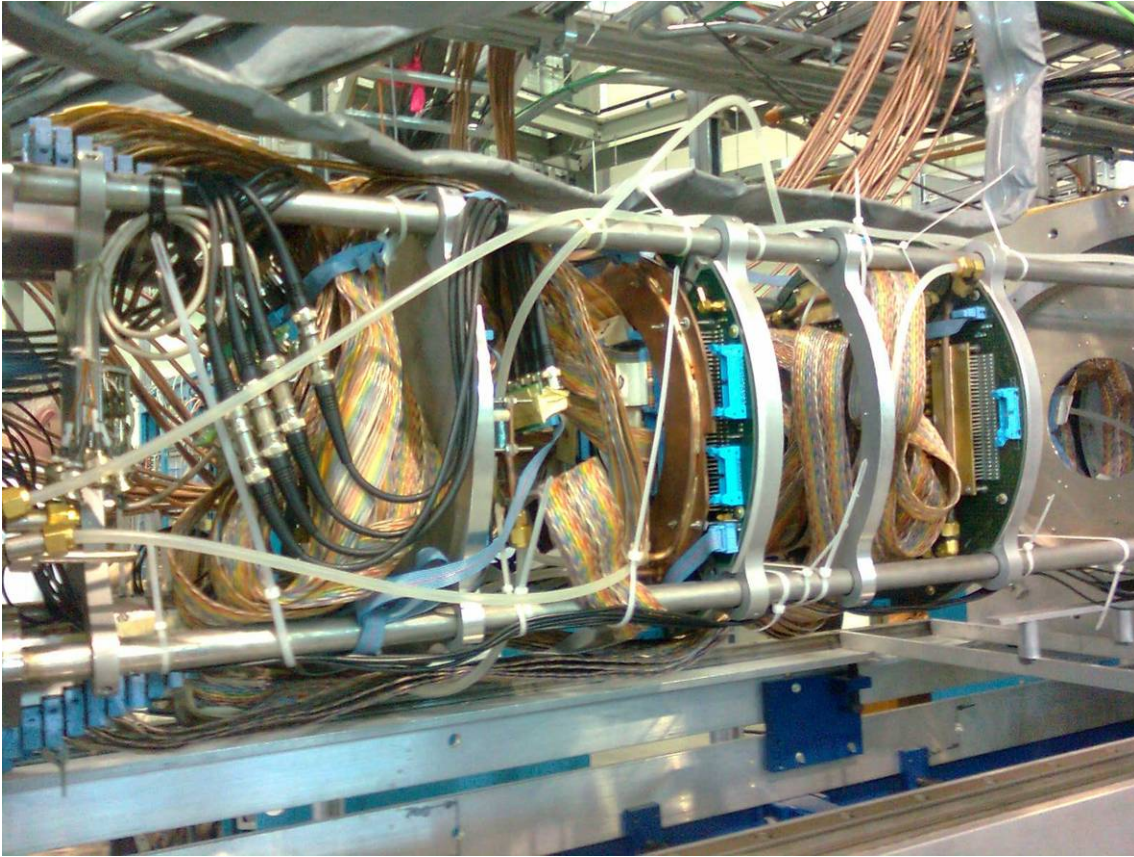
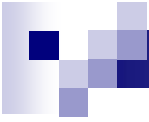


aim 1): measure
 particle
 S2 detector:
 alpha particles
 (thin ta



field for $^{18}\text{Ne} + \alpha$
 by $\Delta E - E$ technique
 CD-PAD:
 heavy ions
 $^{18}\text{Ne}, ^{21}\text{Na}$







Gavin Lotay

Philip Salter

Tom Davinson

work in progress...

S1103 Collaboration

University of Edinburgh

University of York

TRIUMF

Special thanks to Pat Walden, Anffany Chen
TRIUMF Staff: in particular, Bob, Marco, Sophia, Colin
and all the ISAC II operators